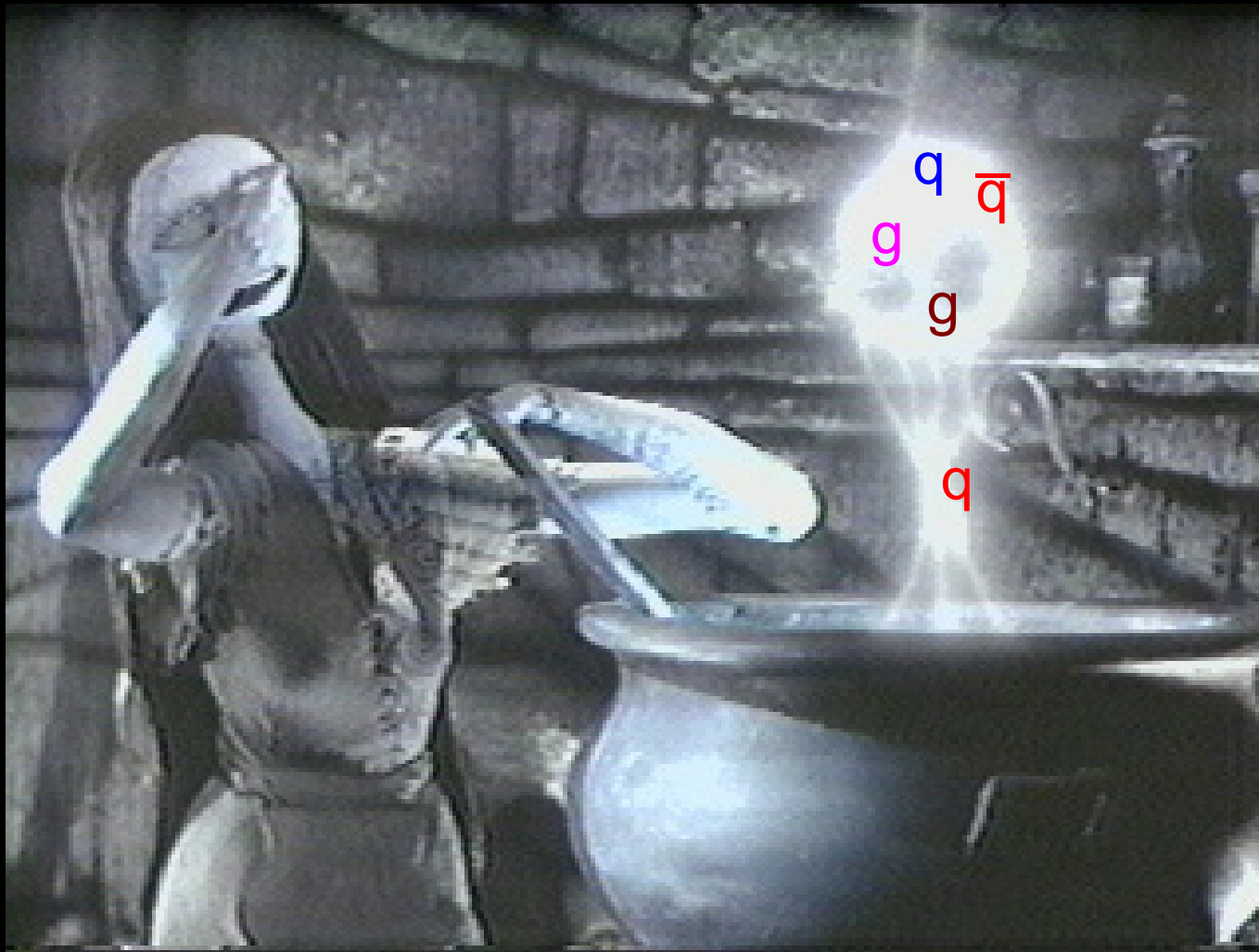


# Weakly Coupled Quark Gluon Plasmas

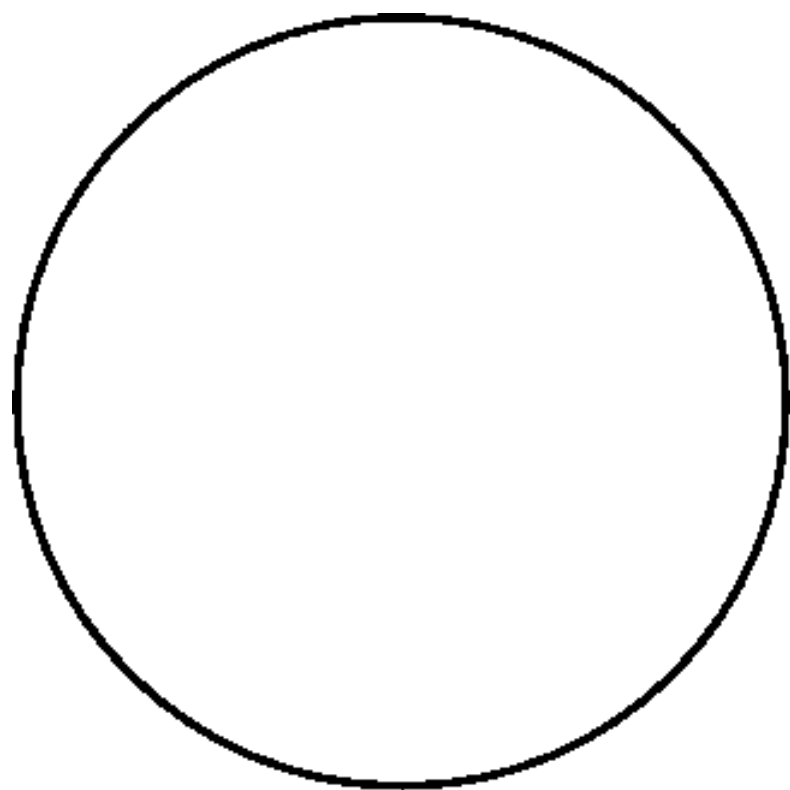


Peter Arnold, University of Virginia

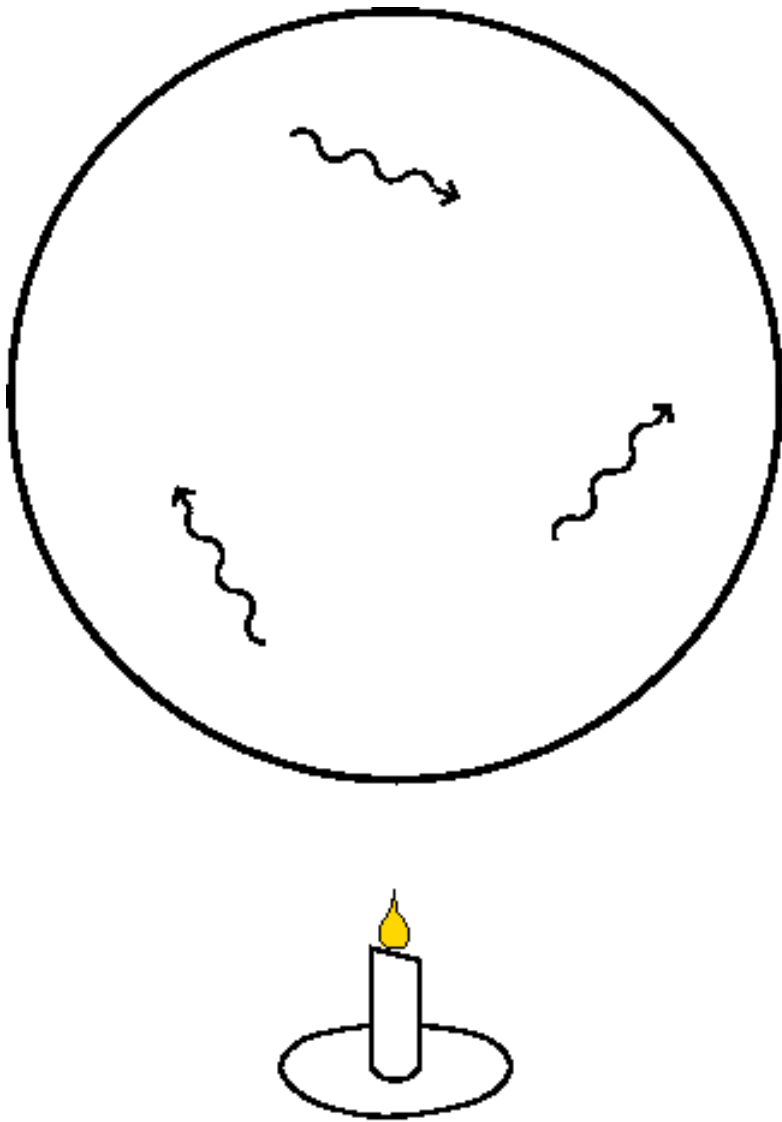
# Notice

This presentation has been carefully purged of anything that might cause embarrassment to the laboratory.

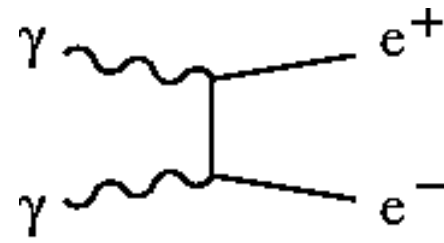
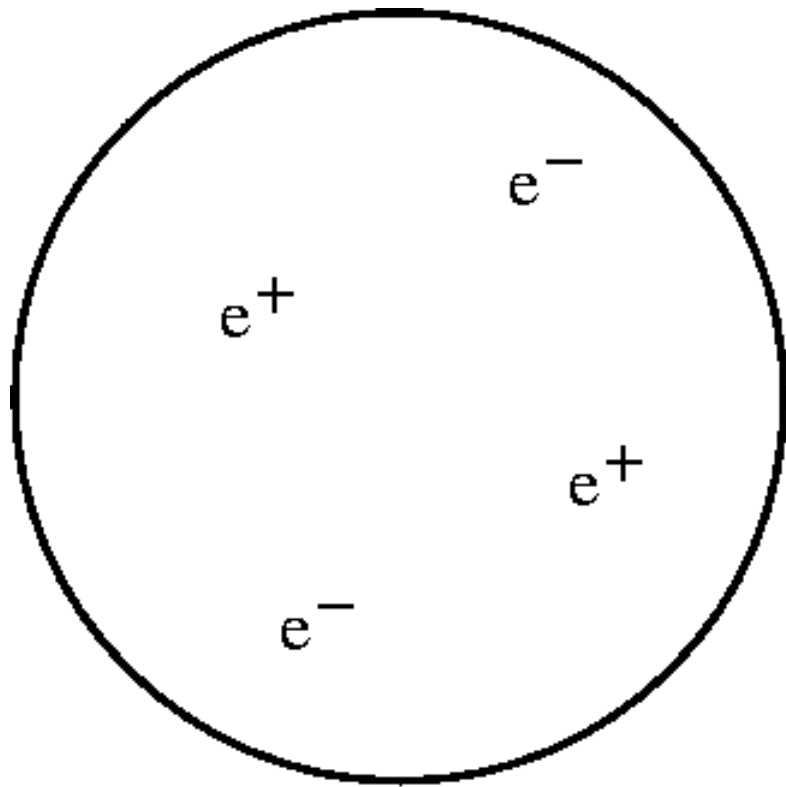




blackbody radiation

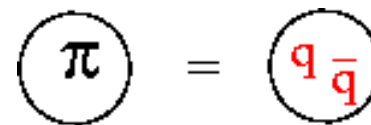
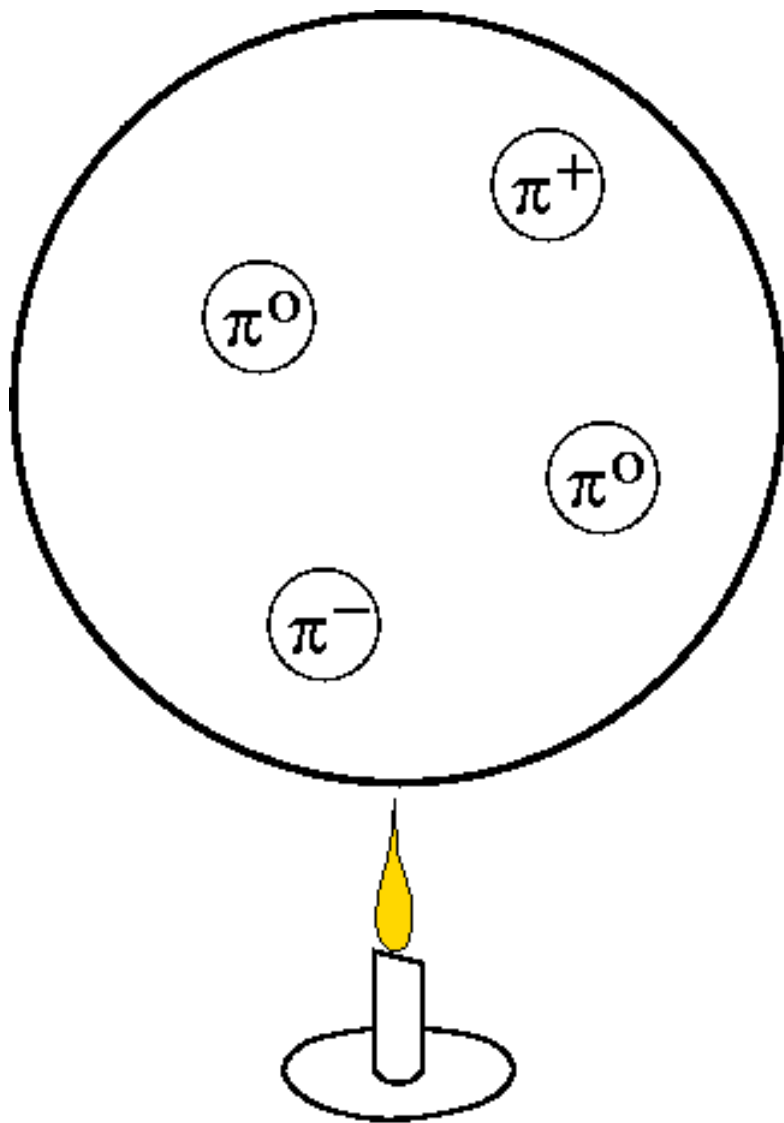


$$k_B T \sim 0.5 \text{ MeV}$$



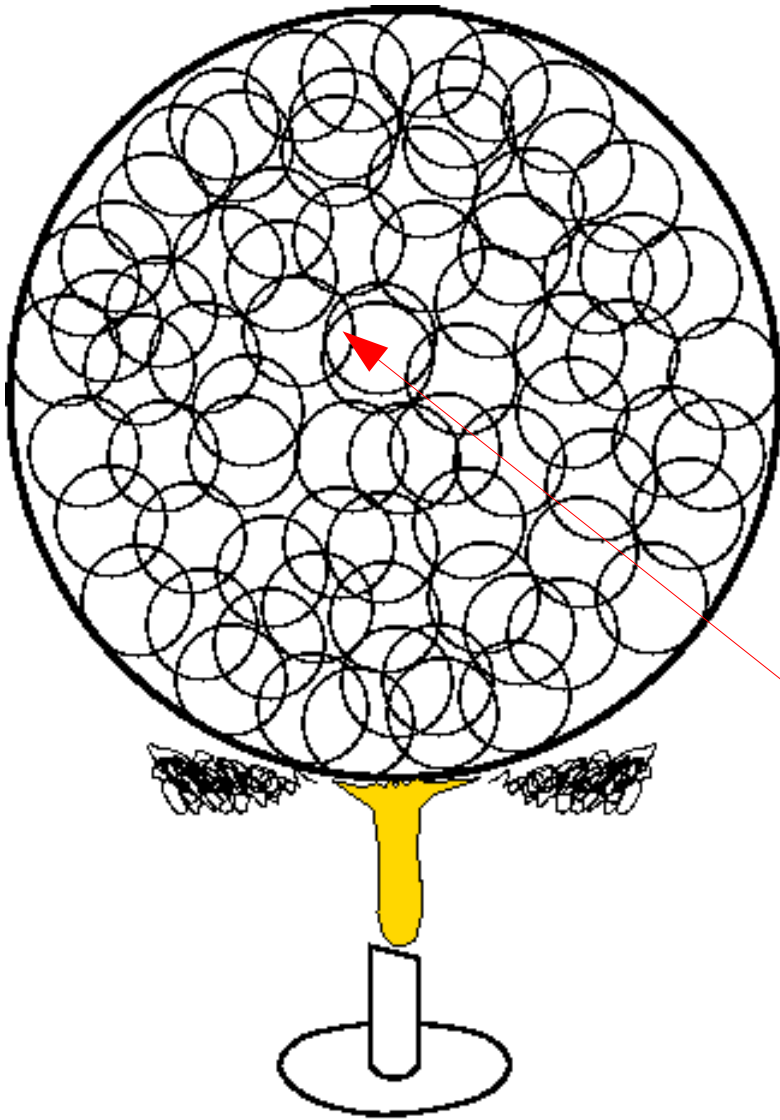
$$\gamma \gamma \rightleftharpoons e^+ e^-$$

$$k_B T \sim 50 \text{ MeV}$$



Higher T  $\longrightarrow$  higher density

$$k_B T \sim 200 \text{ MeV}$$



$$\pi = q \bar{q}$$

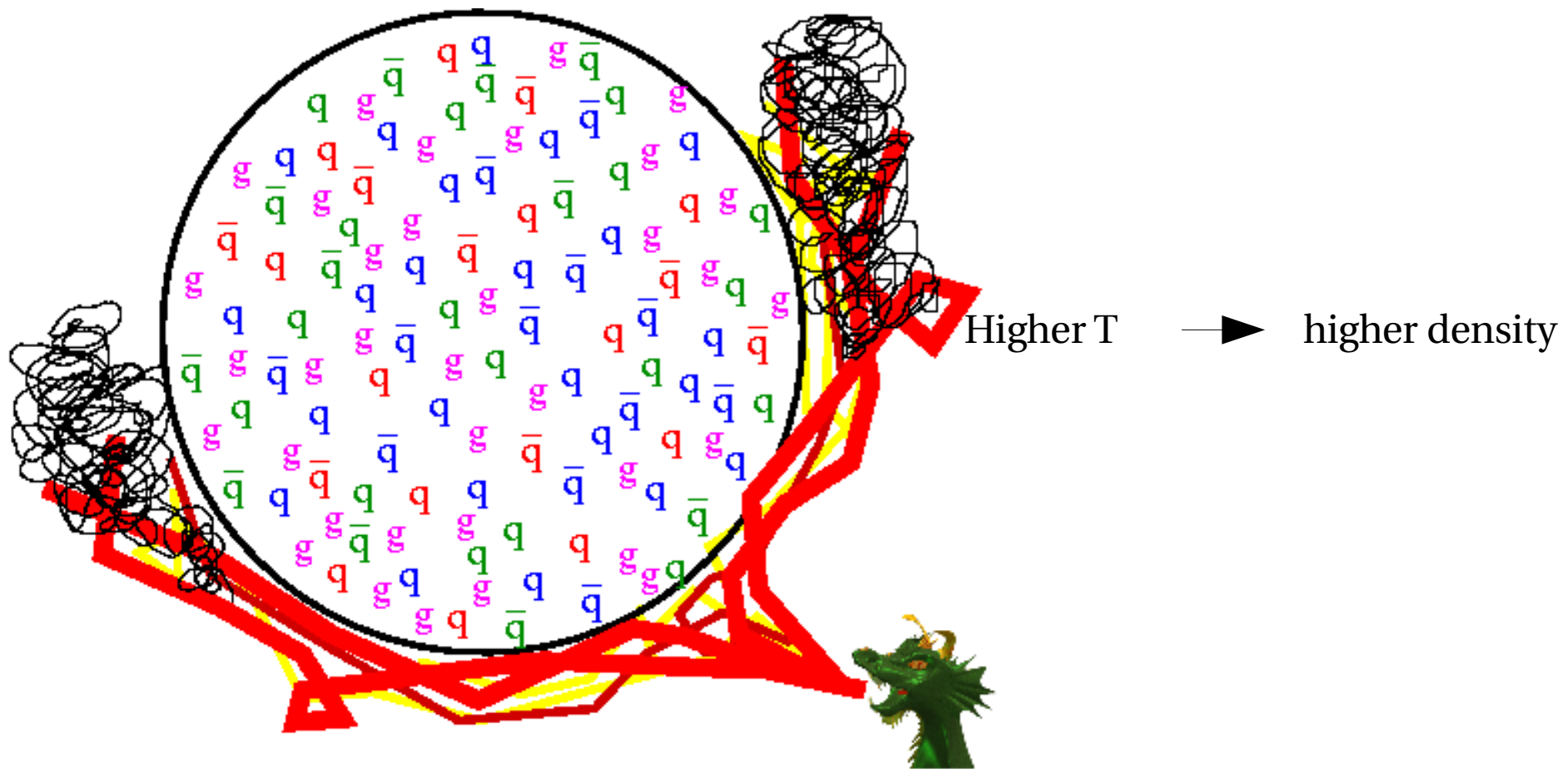
Higher  $T$   $\rightarrow$  higher density

Who do I  
belong to?

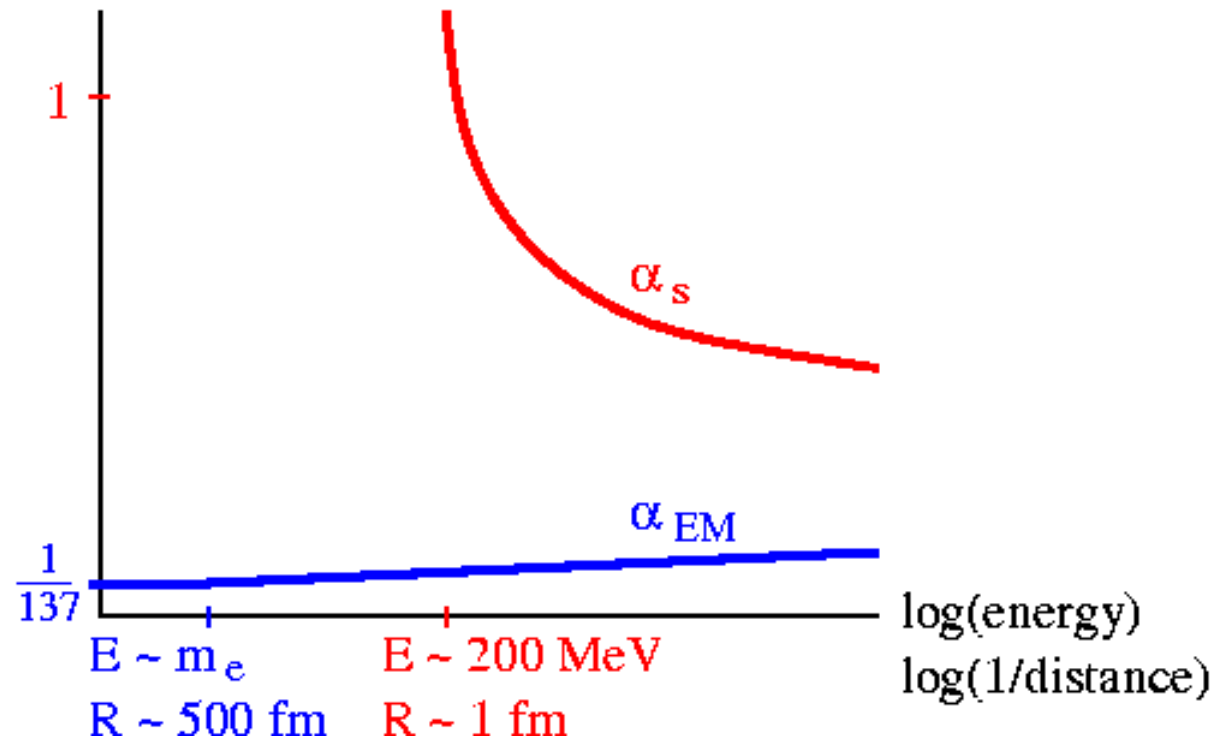
$q$



$$k_B T \gg 200 \text{ MeV}$$



# Also: Asymptotic Freedom



$$\alpha \equiv \frac{e^2}{4\pi\hbar c}$$

Higher temperature  $\rightarrow$  smaller coupling  $\alpha_s$

# Why bother with weak coupling?

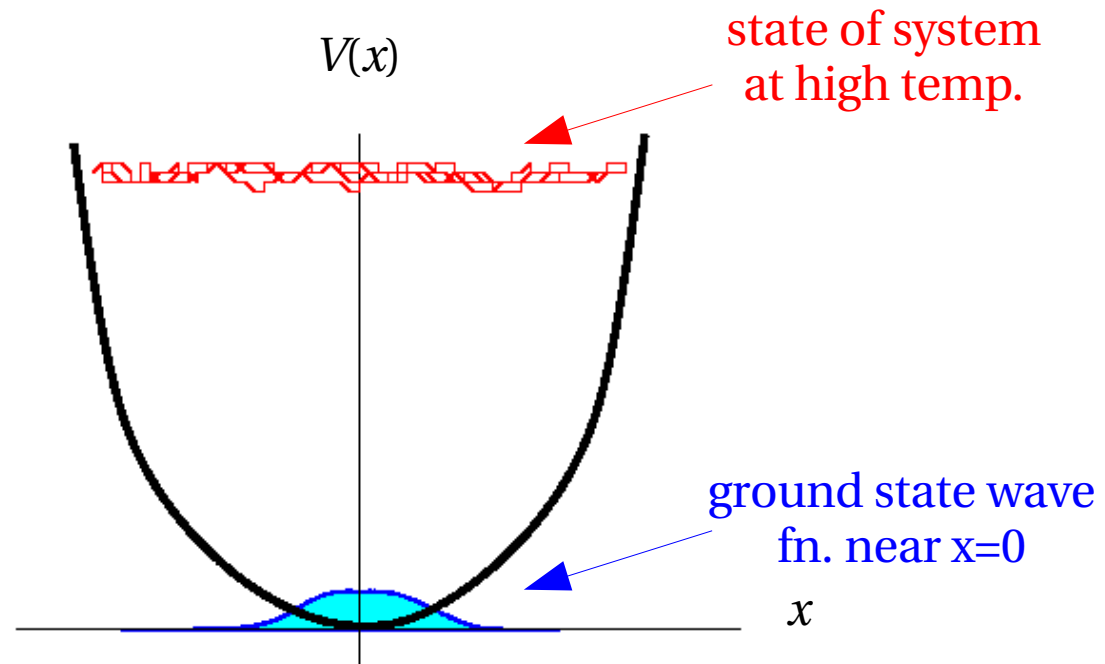
It's one of the few limits where we can do calculations from first principles.

- Lattice simulations  
imaginary time: difficult to apply to real-time response.
- Consider  $T$  very large so that running coupling  $\alpha_s(T)$  is small.
- Change the theory (add lots of supersymmetry, take # colors to infinity) and then use AdS/CFT methods to study limit of *really* big coupling.

# Isn't weak coupling easy?

## Counter-example

$$V(x) \sim \omega_0^2 x^2 + g^2 x^4$$



Note 1: problems with perturbation theory if  $T$  high enough.

Note 2: For fixed  $T$ ,  $\omega_0 \rightarrow 0 \implies$  non-perturbative.

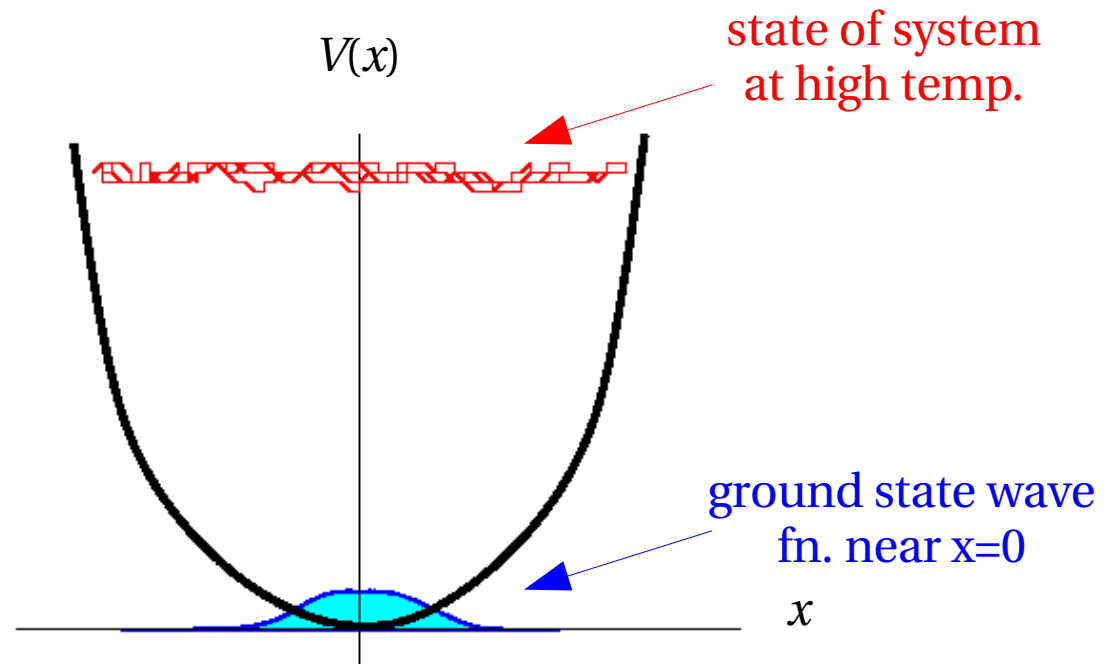
For gauge theory,  $\omega_k \sim k \rightarrow 0 \implies$  non-perturbative.

**Moral: small coupling expansion not the same as the perturbative expansion.**

# Isn't weak coupling easy?

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*Moral:* small coupling expansion not the same as the perturbative expansion.

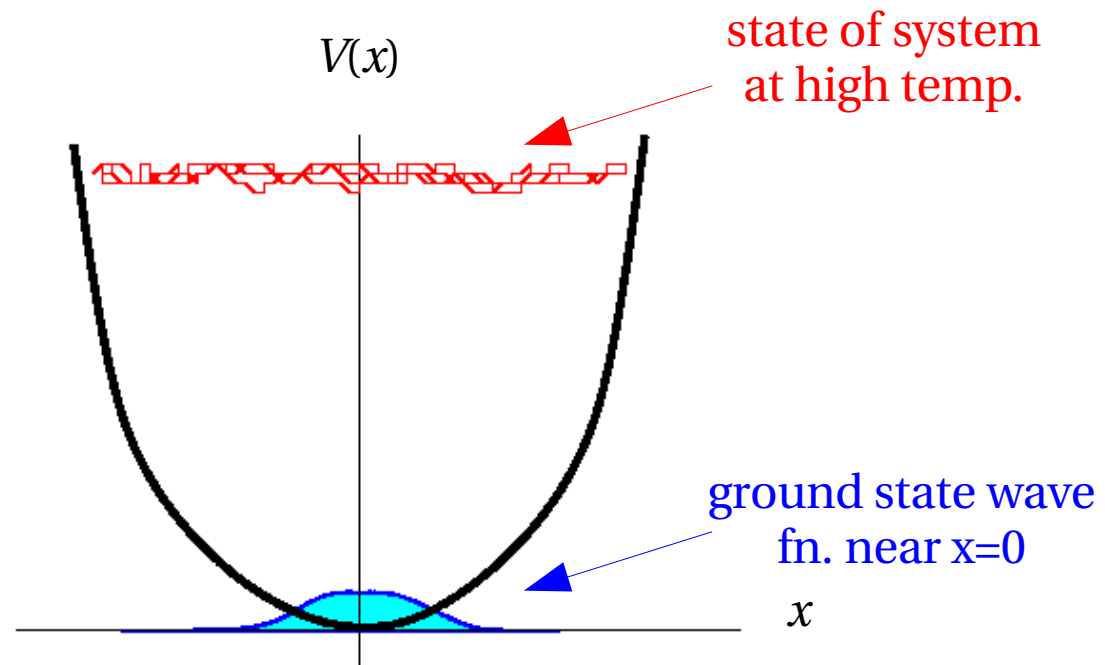
**Example:**  $P = \#T^4 [1 + \#g^2 + \#g^3 + g^4(\# \ln g + \#) + \#g^5 + g^6(\# \ln g + \#) + \dots]$

non-perturbative ↗

# Isn't weak coupling easy?

## Counter-example

$$V(x) \sim \omega_0^2 x^2 + g^2 x^4$$



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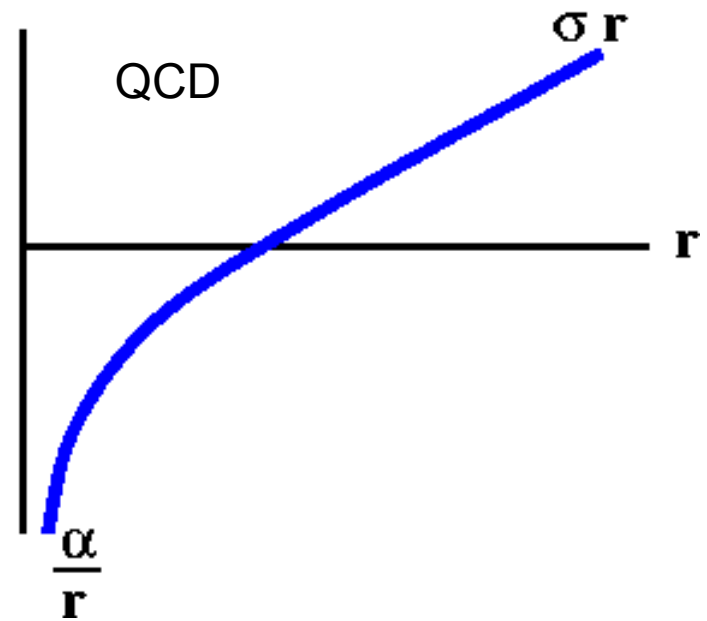
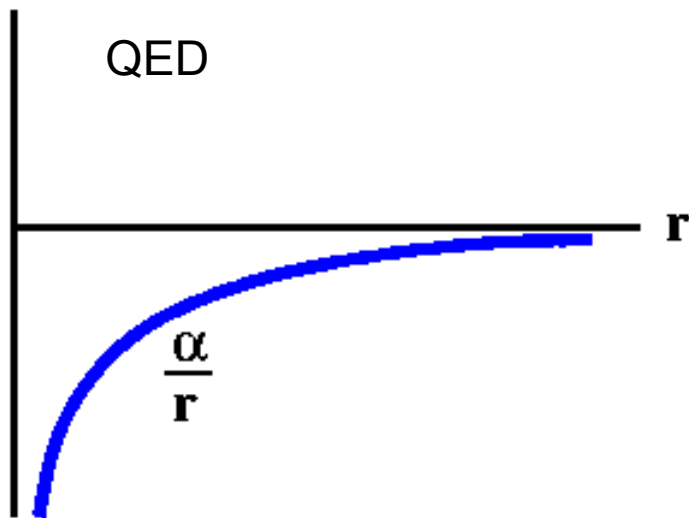
**Example:**  $P = \#T^4 [1 + \#g^2 + \#g^3 + g^4(\# \ln g + \#) + \#g^5 + g^6(\# \ln g + \#) + \dots]$

units:  $\hbar=c=k_B=1$

non-perturbative

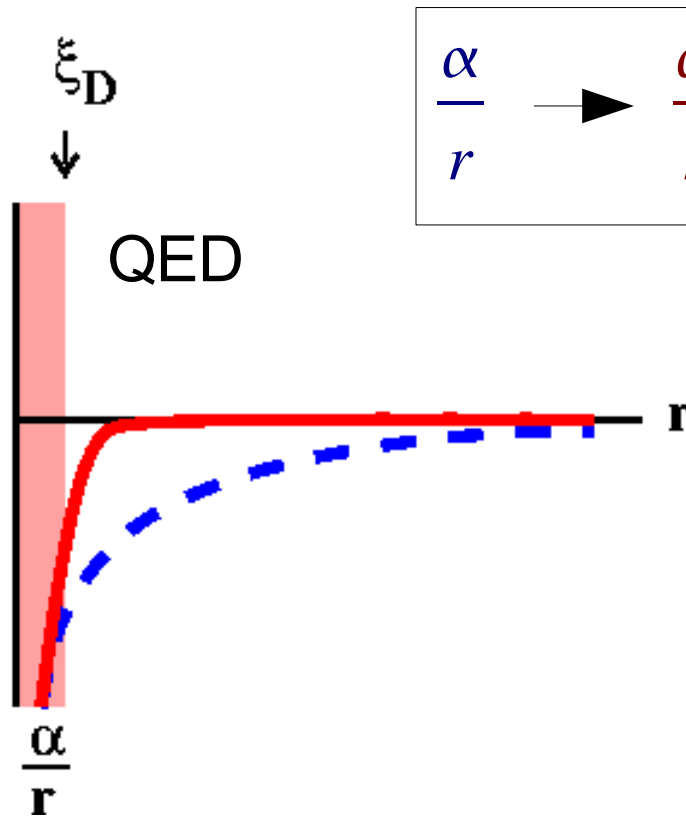
# Deconfinement as Debye Screening

Potential energy between 2 charges in vacuum

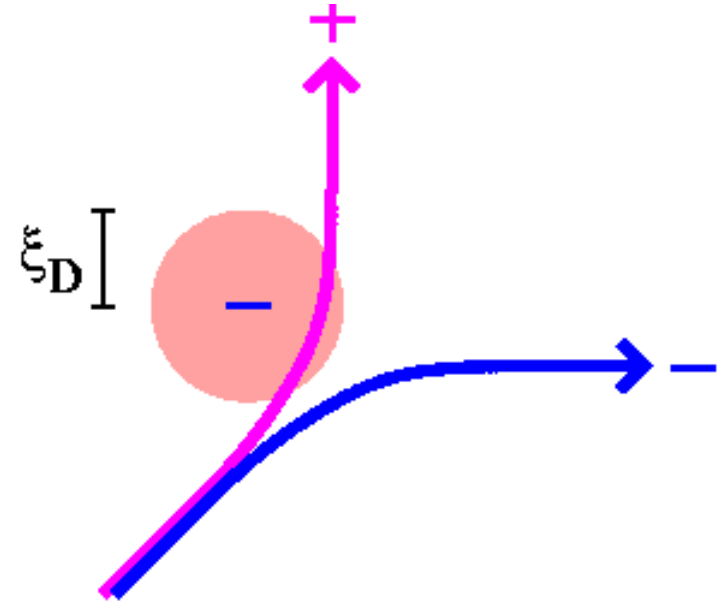


# Deconfinement as Debye Screening

In a medium with free charges:



$$\frac{\alpha}{r} \longrightarrow \frac{\alpha}{r} e^{-r/\xi_D}$$



Debye screening length:

$$\xi_D \sim \left( \frac{T}{e^2 n} \right)^{1/2}$$

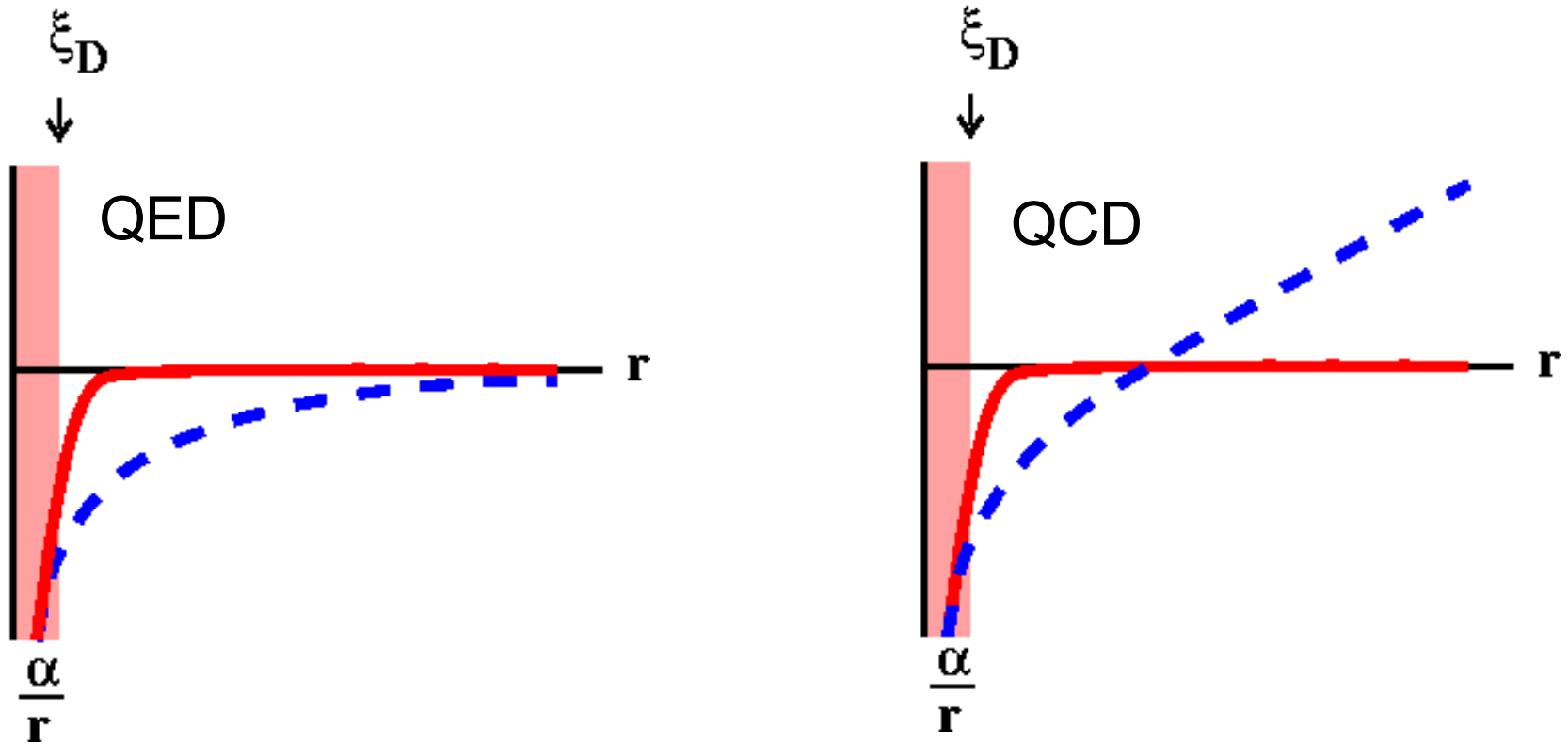
ultra-relativistic:  $n \sim T^3 \longrightarrow \xi_D \sim \frac{1}{eT}$

Higher temperature  $\longrightarrow$  smaller Debye radius



# Deconfinement as Debye Screening

In a medium with free charges:



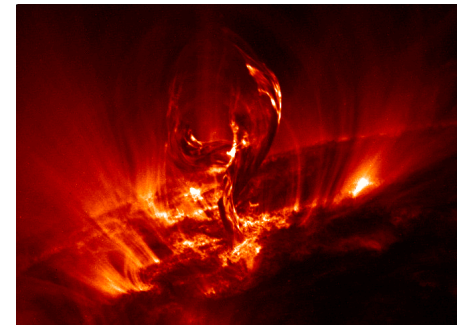
Higher temperature  $\rightarrow$  smaller Debye radius

The Debye effect screens electric fields. In contrast:

Magnetic fields are not screened in a plasma.

So

QED: magnetic forces are still long range



QCD: could there be confinement of colored currents?

—► no long range colored B fields?

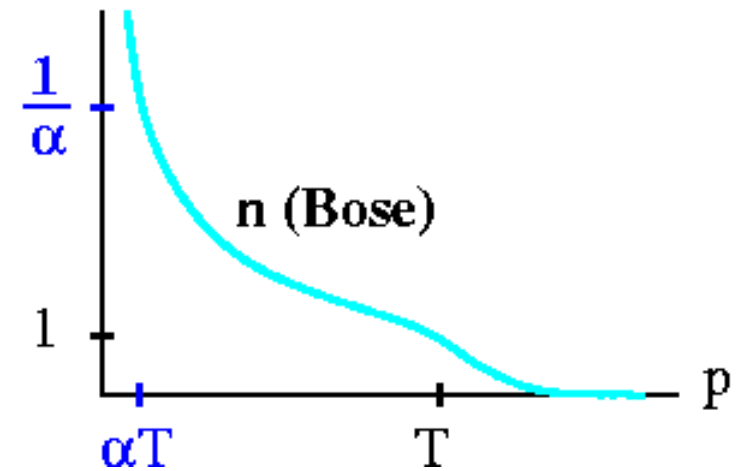
Version for particle theorists: Do spatial Wilson loops still have area-law behavior?

YES, and at very short distances too!

$$n_{\text{Bose}} = \frac{1}{e^{\beta E} - 1} \rightarrow \frac{T}{E} \quad \text{as } E \rightarrow 0$$

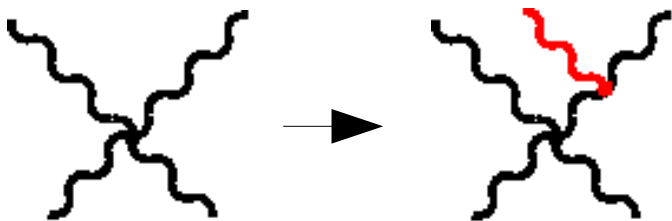
For massless bosons,

$$E \sim p \sim \alpha T \quad \rightarrow \quad n_{\text{Bose}} \sim \frac{1}{\alpha}$$



Photons don't directly interact with each other, but gluons do.

**Result:** Perturbation theory breaks down for gluons with  $p \sim \alpha T$ .



costs  $|g|^2 \sim \alpha$  for extra interaction

$n_{\text{Bose}} \sim \frac{1}{\alpha}$  for density of extra gluons

---

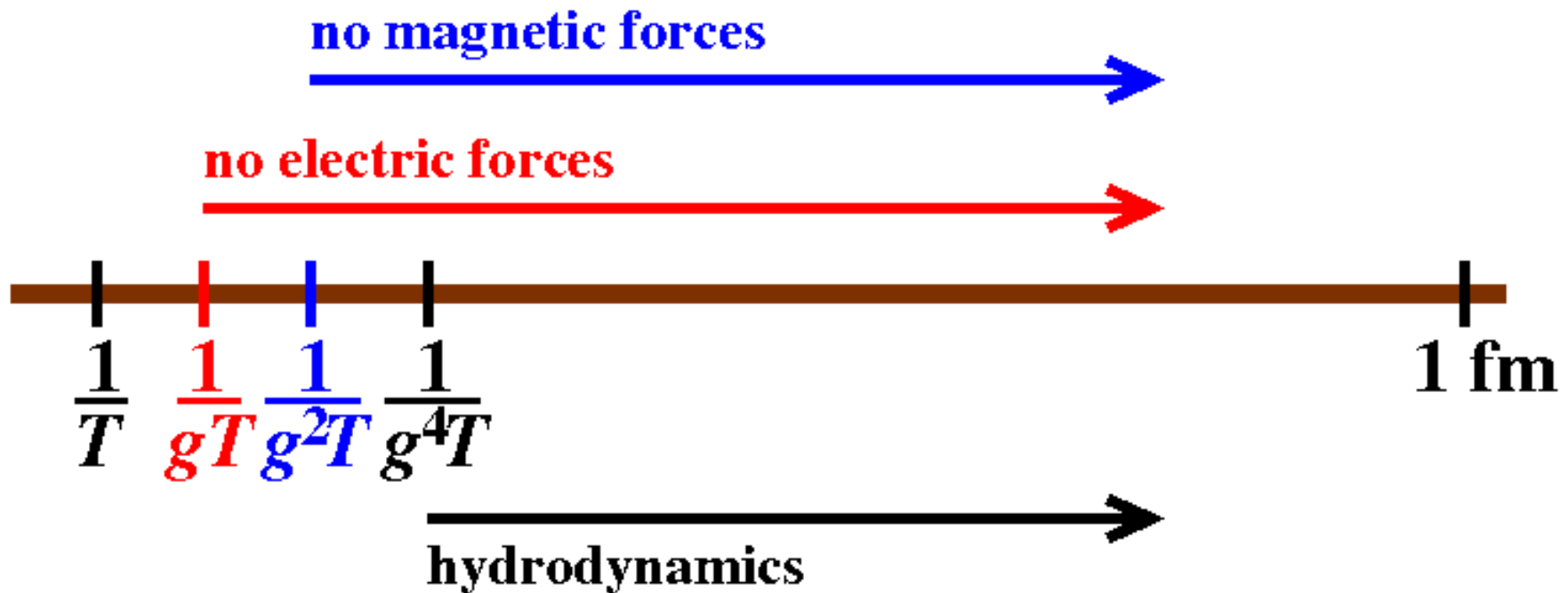
1 total

# Summary

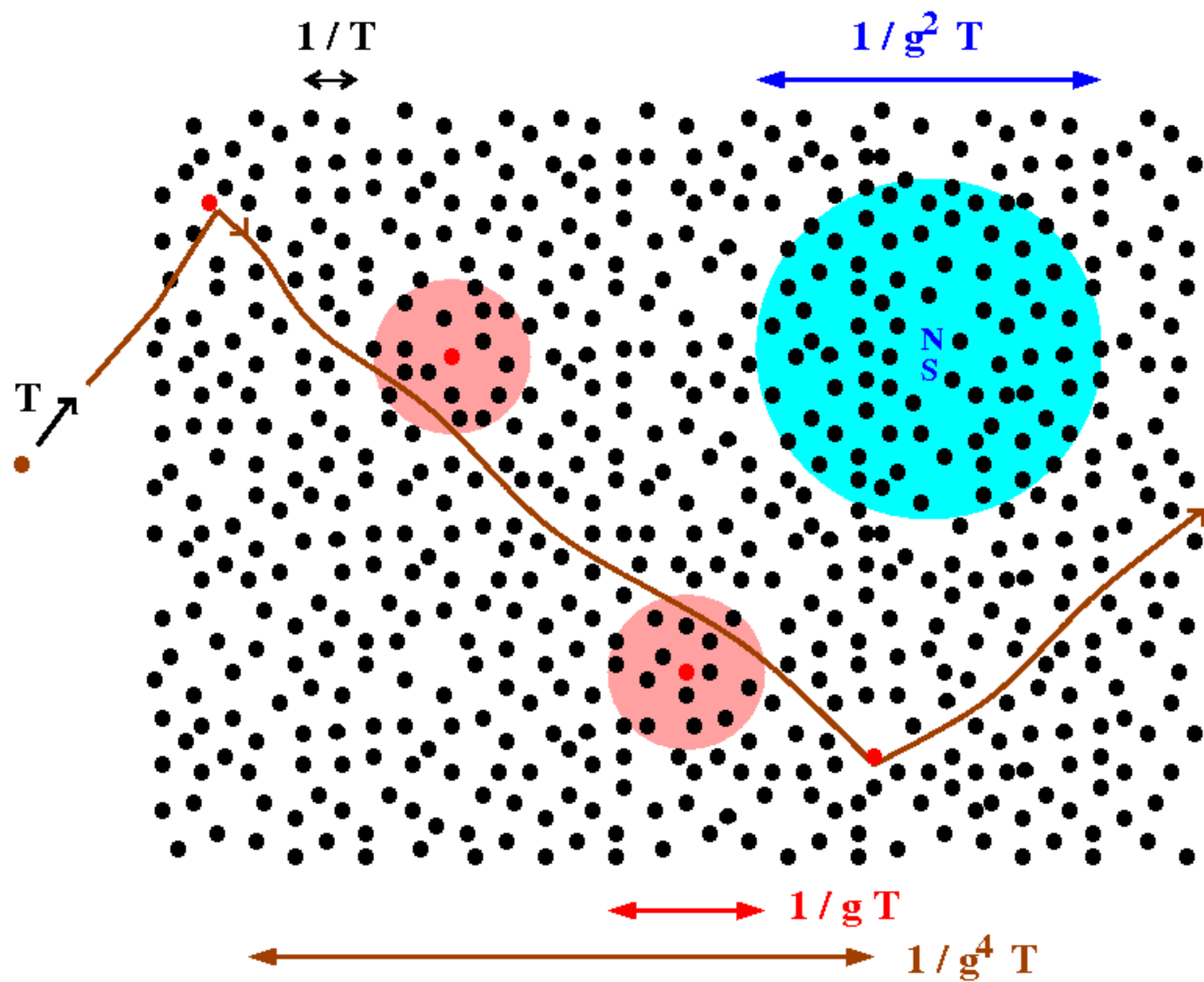
Note: “ $g$ ” is QCD analog of “ $e$ ”

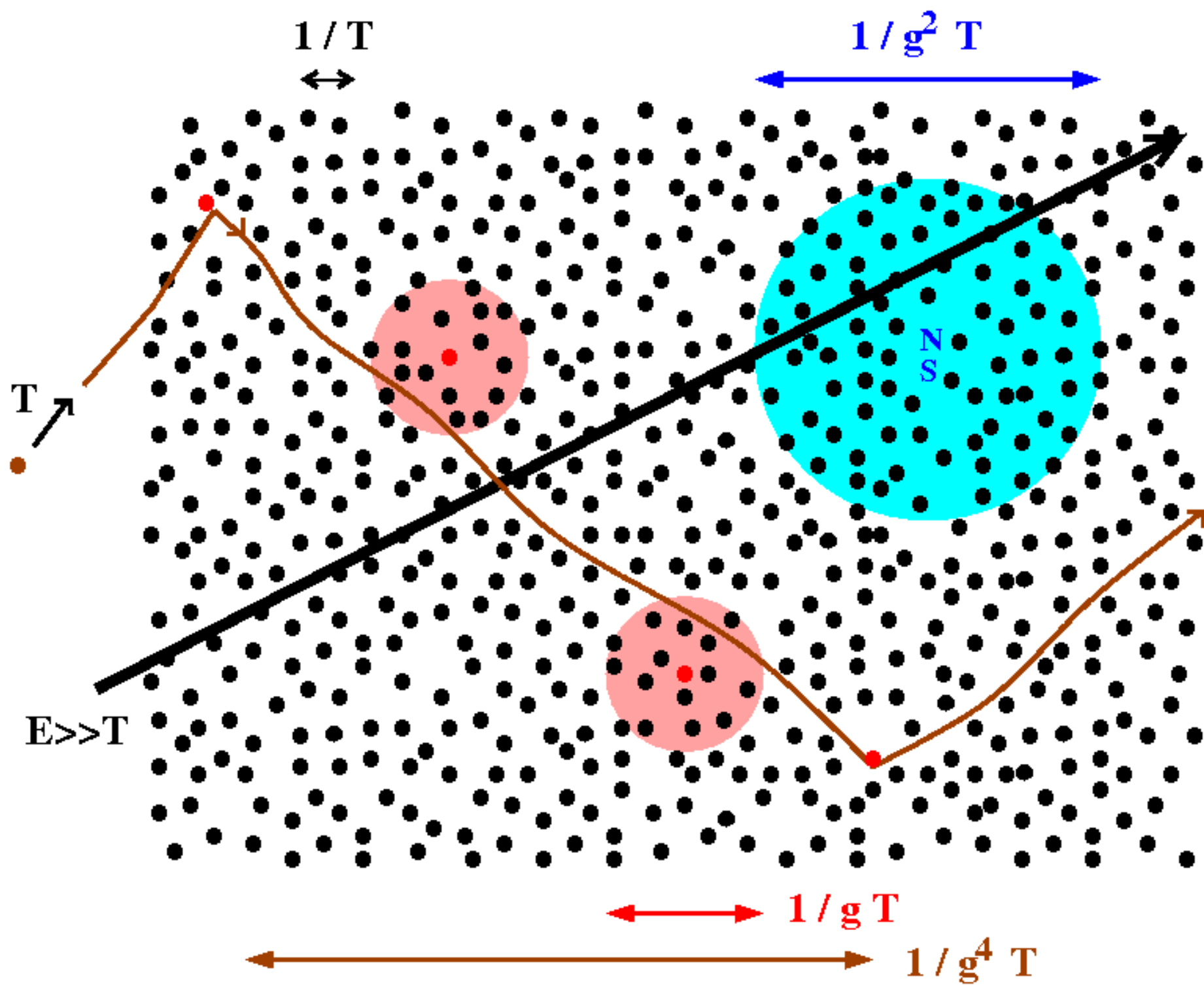
electric screening at  $\xi_D \sim \frac{1}{gT}$   $\rightarrow$  no charge confinement

no traditional magnetic screening  $\rightarrow$  current confinement at  $\frac{1}{g^2 T}$



Long distance physics is hydrodynamics,  
not colored MHD.





# Landau-Pomeranchuk-Migdal (LPM) effect

## What is the LPM Effect?

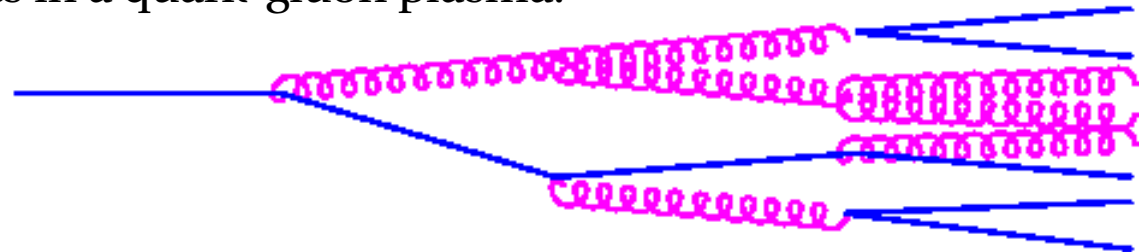
A coherence effect that complicates calculations of bremsstrahlung or pair production when a very high energy particle scatters from a medium.

## Places it comes up in QED

- Very high energy cosmic rays showering in the atmosphere.
- Certain beam dump experiments designed to measure the LPM effect.

## Places it comes up in QCD

- Energy loss of high energy jets in a quark-gluon plasma.



- Complete leading-order calculations of the viscosity and other transport coefficients of a weakly-coupled quark-gluon plasma.



# The LPM Effect

*Naively*

ely

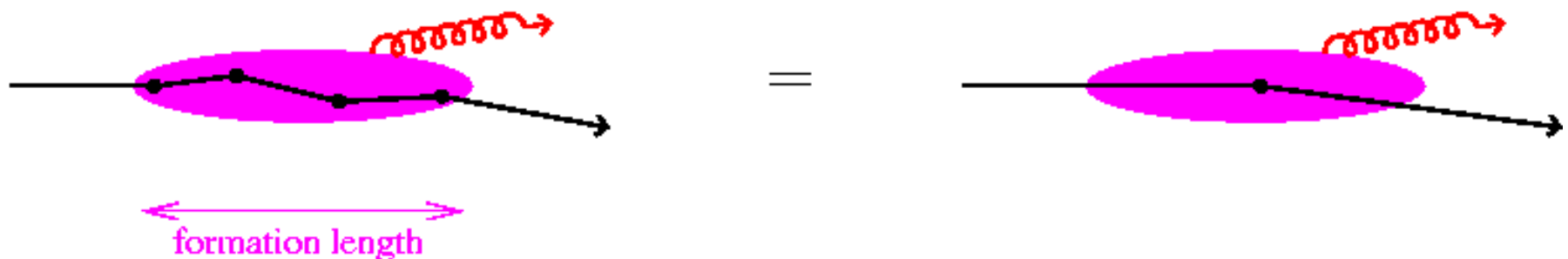
brem rate  $\sim n\sigma v \sim (\text{density of scatterers}) \times \left| \begin{array}{c} \text{E} \\ \downarrow \gamma \end{array} \right|^2 \times 1$

### **Problem**

At very high energy,

probabilities of brem from successive scatterings no longer independent;

brem from several successive (small angle) collisions not very different from brem from one collision.



formation length  $\propto \sqrt{E}$

*Result:* a reduction of the naive brem rate.



# Example: stopping distance (in a infinite medium)

If LPM effect ignored:

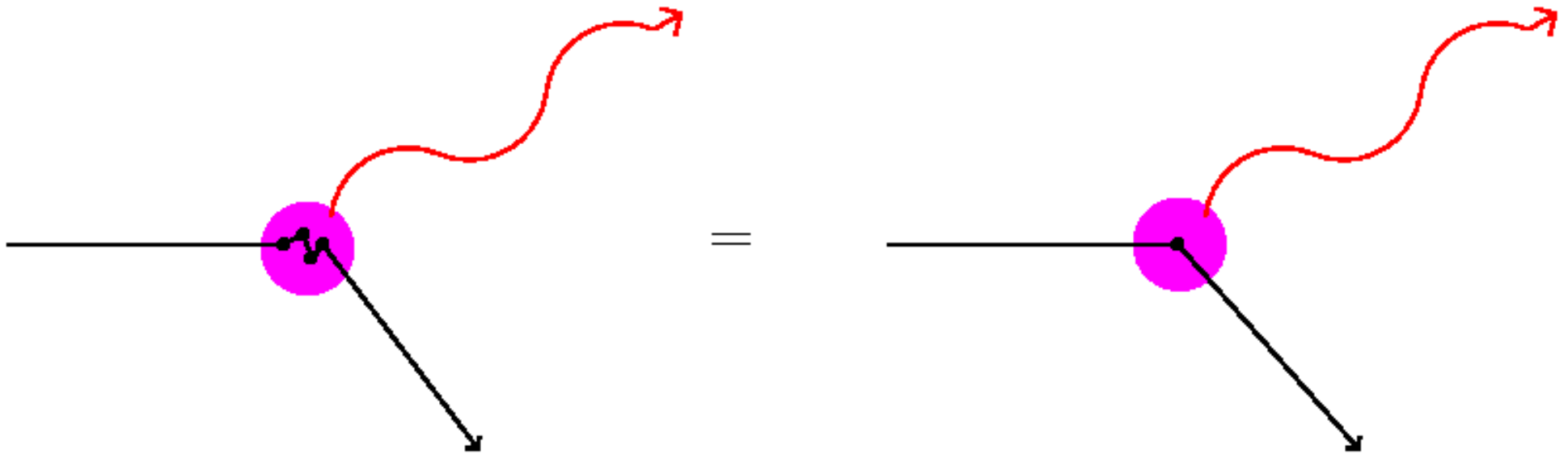
$$\text{stopping distance} \propto \ln E$$

Actual result (weak coupling):

$$\text{stopping distance} \propto \left( \frac{E}{\ln E} \right)^{1/2}$$

# The LPM Effect (QED)

Warm-up: Recall that light cannot resolve details smaller than its wavelength.

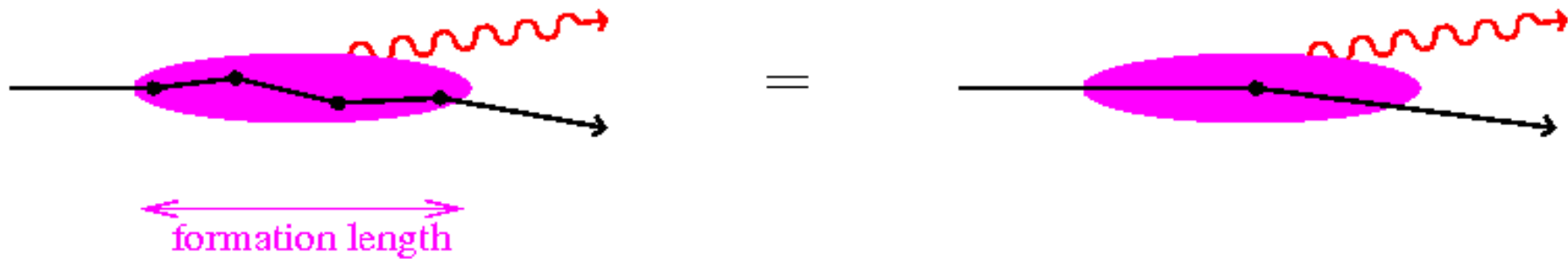


[Photon emission from different scatterings have same phase  $\rightarrow$  coherent.]

*Now: Just Lorentz boost above picture by a lot!*

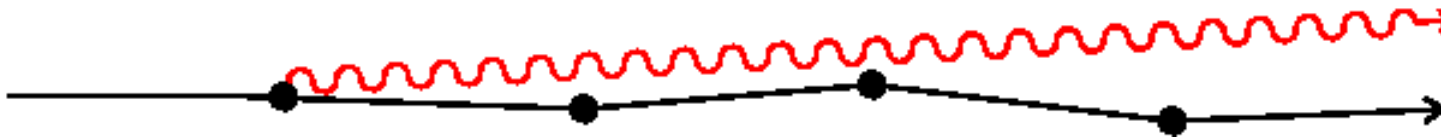


# The LPM Effect (QED)

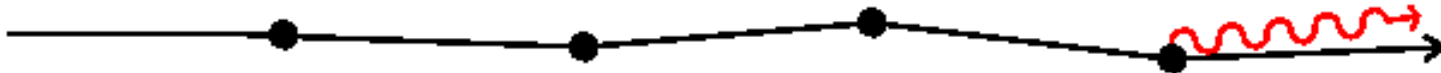


- Note:
- (1) **bigger  $E$**  requires bigger boost  $\rightarrow$  more time dilation  $\rightarrow$  **longer formation length**
  - (2) big boost  $\rightarrow$  this process is **very collinear**.

# An alternative picture



*versus*

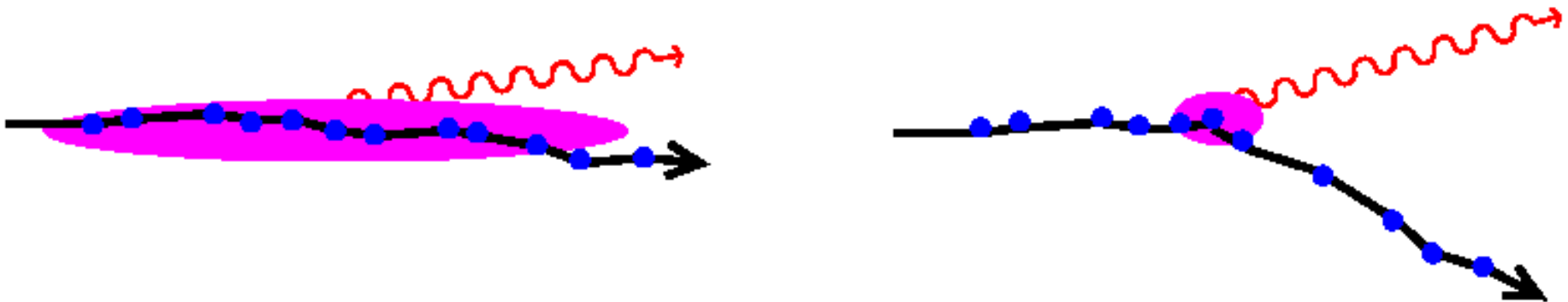


Are these two possibilities in phase? Or does the interference average to zero?

IN PHASE	if (i) everything is nearly collinear	✓
	(ii) particle and photon have nearly same velocity	✓ ( <i>speed of light</i> )

The important point:

The more collinear the underlying scattering, the longer the formation time.



*Note:* the formation length

***depends on*** the net angular deflection during the formation length, which  
***depends on*** the formation length

[ Self-consistency  $\rightarrow$  standard parametric formulas for formation length. ]

# The LPM Effect (QCD)

There is a qualitative difference for *soft* bremsstrahlung.:

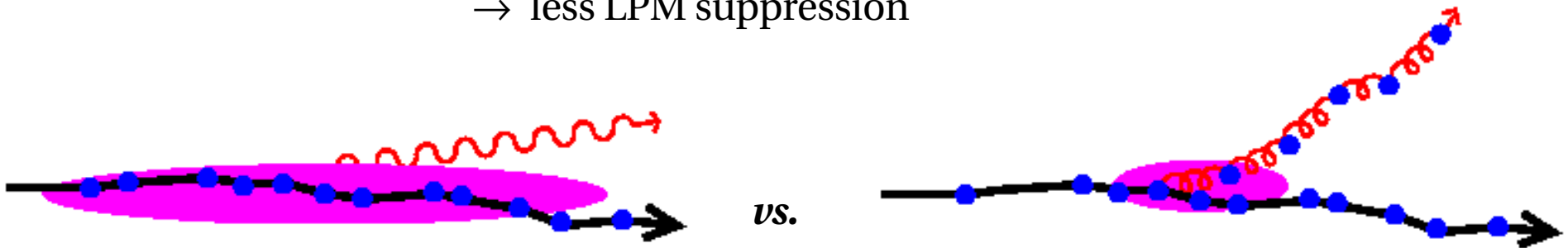
## QED

Softer brem photon → longer wavelength  
→ less resolution  
→ more LPM suppression

## QCD

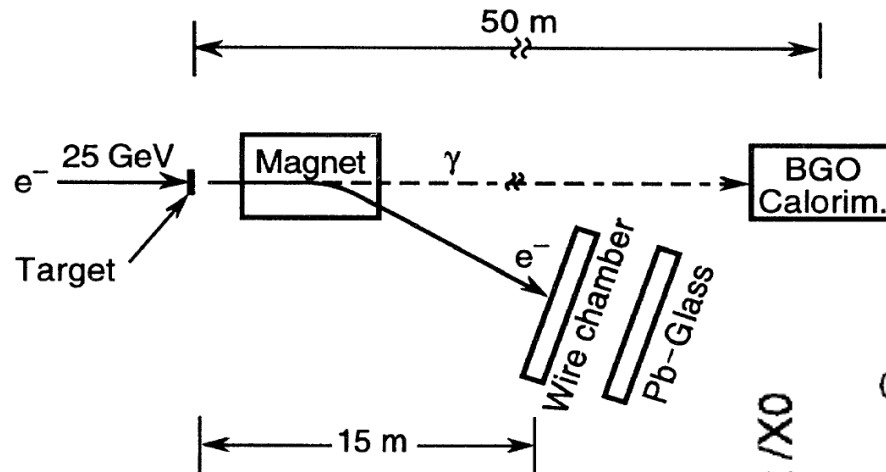
Unlike a brem photon, a brem gluon can easily scatter from the medium.

Softer brem gluon → easier for brem gluon to scatter  
→ less collinearity  
→ less LPM suppression



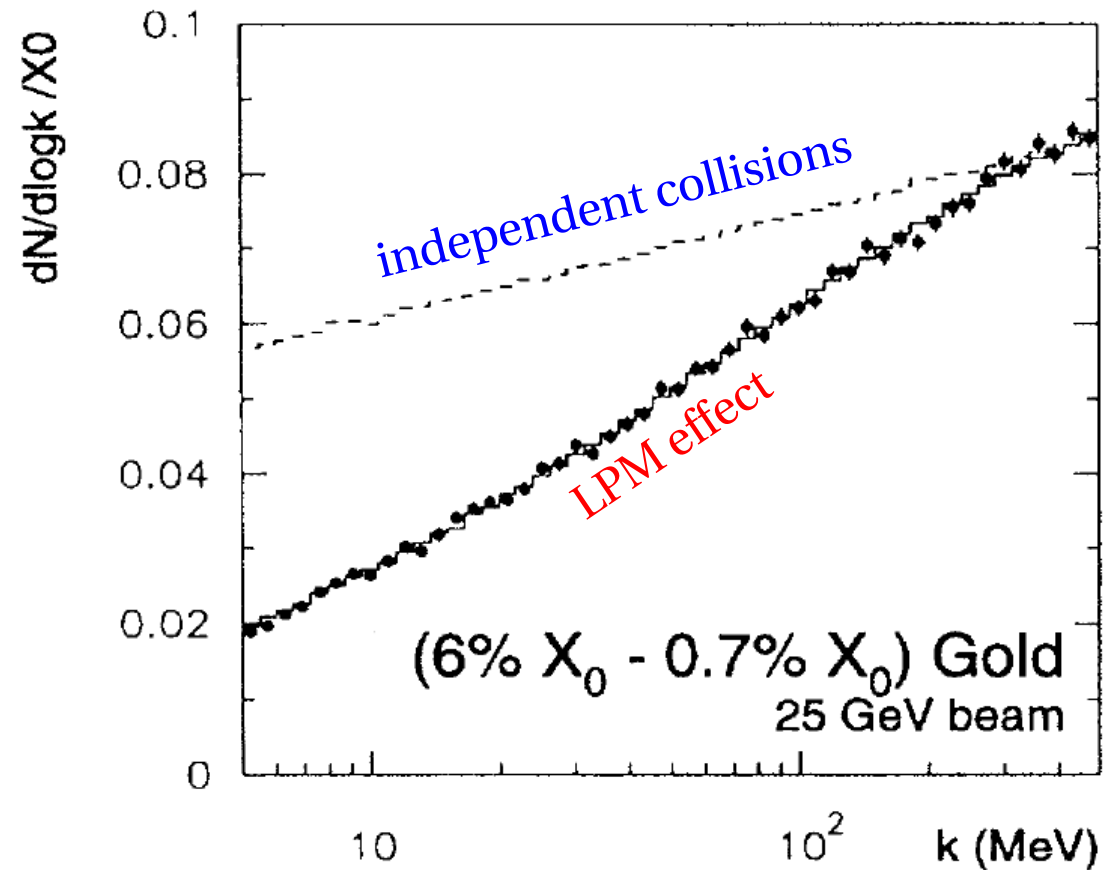
**Upshot:** Soft brem more important in QCD than in QED (for high- $E$  particles in a medium)

# Experimental Measurement of LPM (QED)

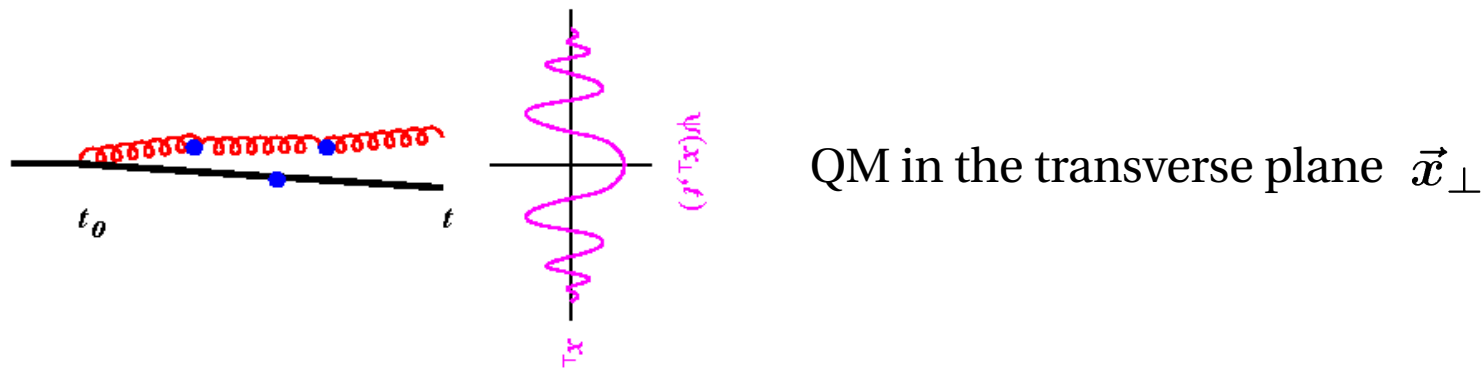


SLAC E-146

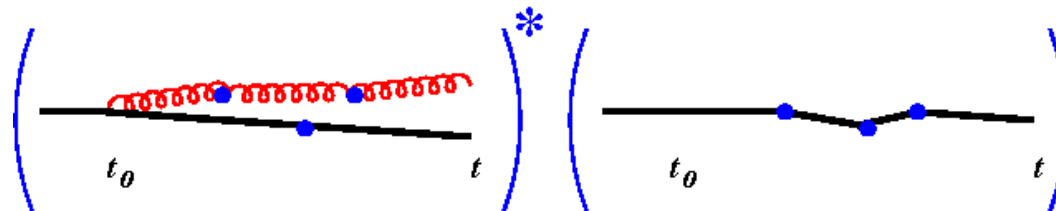
Phys. Rev. Lett. **75** (1995) 2949.



# 2-dimensional Quantum Mechanics



But what we really need is the time evolution of the interference



[QED: Migdal '56]

random-averaged over the locations and types of scatterers in the plasma.  
Evolution of this interference is described by a 2-dimensional Schrödinger eq. with

$$H(t) = \frac{p_\perp^2}{2\mathcal{M}} - i\Gamma(x_\perp, t)$$

$\uparrow$   $\uparrow$   $\uparrow$   
 $E_x(1-x)$  for non-uniform media  
 non-Hermitian

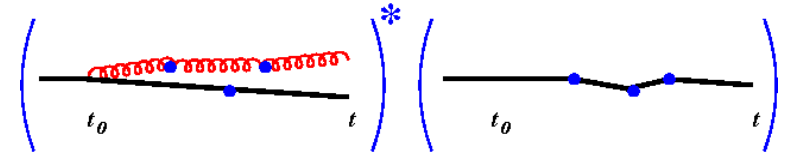
What assumptions have been made?

screening length  $\ll$  mean free path

$$\frac{\#}{gT} \ll \frac{\#}{g^2T}$$

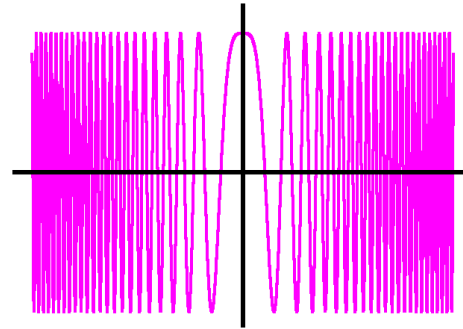
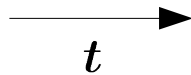


$$H(t) = \frac{p_{\perp}^2}{2\mathcal{M}} - i\Gamma(x_{\perp}, t)$$



Numerically a bit tricky

$\delta(x_{\perp})$  at  $t_0$



for  $\Gamma=0$ , for example

Harmonic Oscillator approximation

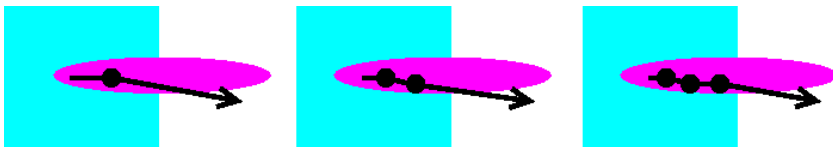
$$H(t) = \frac{p_{\perp}^2}{2\mathcal{M}} - i\hat{q}x_{\perp}^2$$

Turns out to apply to thick media

at very large energies:  $\ln(E/T) \gg 1$

QM Perturbation theory (the opacity expansion)

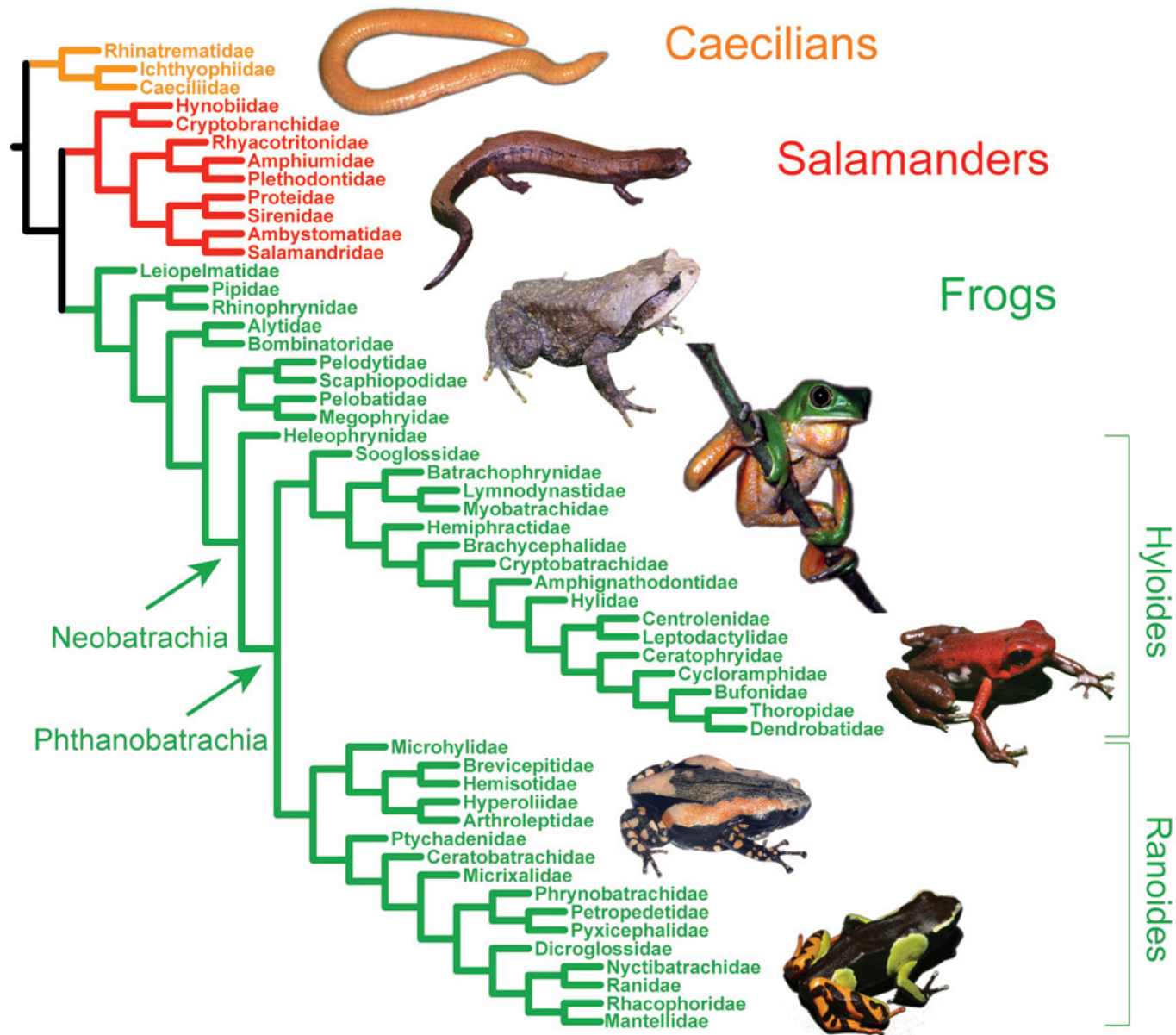
$$H_0 + \delta H(t) = \frac{p_{\perp}^2}{2\mathcal{M}} - i\Gamma(x_{\perp}, t)$$



Applies to thin media

(but needn't be as thin as you might think)

# Taxonomy of Jet Quenching Formalisms



BDMPS-Z/ASW

# Taxonomy of Jet Quenching Formalisms

But I'm going to restrict attention to

- effectively massless partons (no heavy quark jets)
- methods based on the preceding formalism

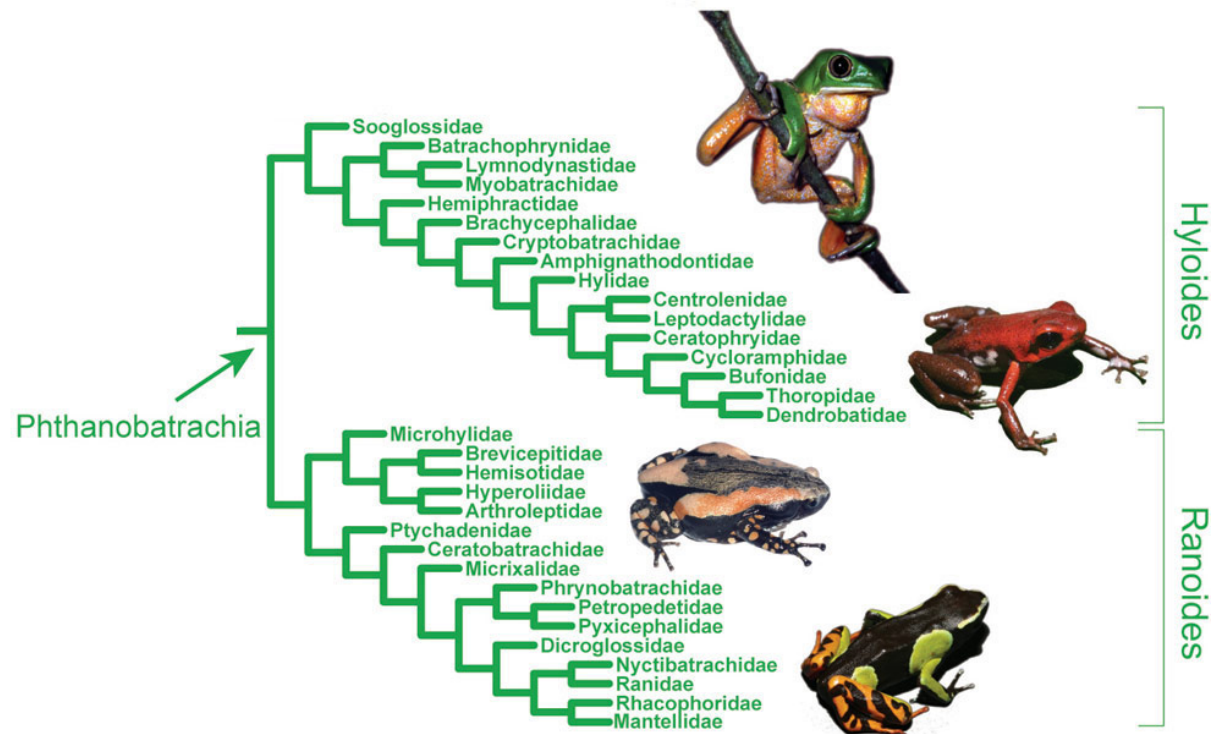
*Apologies in particular to the “higher twist” (HT) jet quenching members of the community.*

# Taxonomy of Jet Quenching Formalisms

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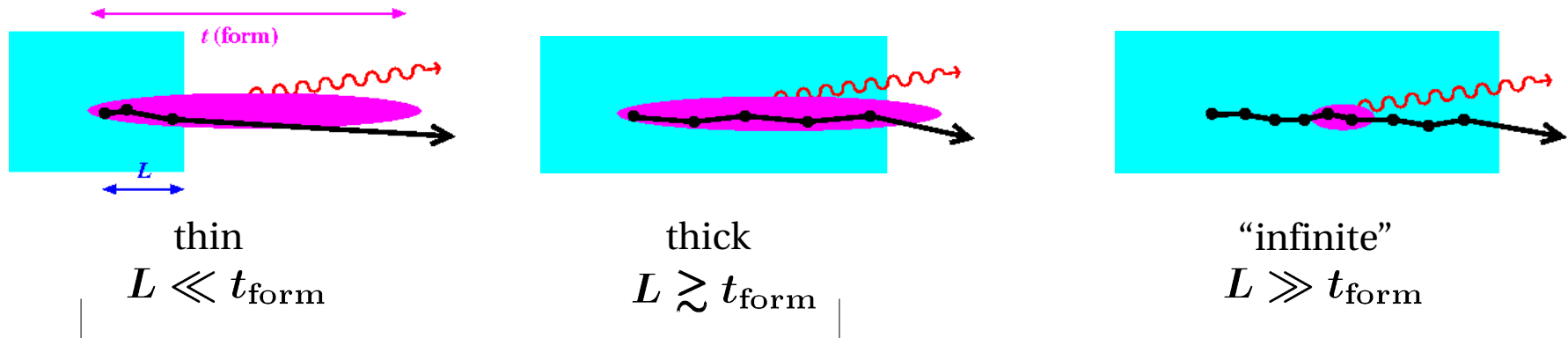
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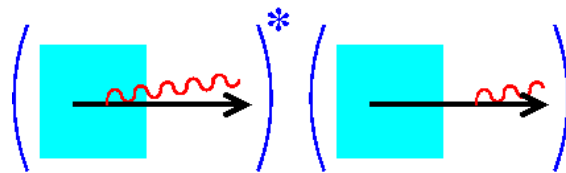


# Taxonomy of Jet Quenching Formalisms

Does it handle thick or thin media, or both?

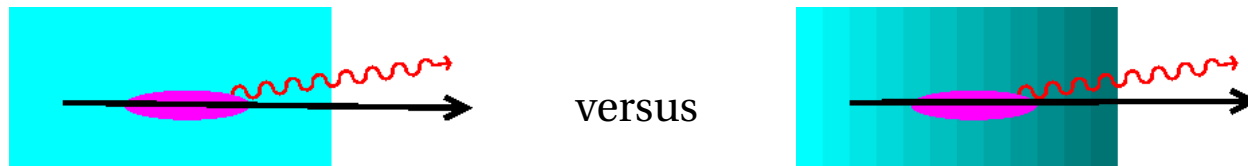


Must deal with medium-vacuum interference:



N/A

Does it handle non-uniform, time-dependent media?



versus

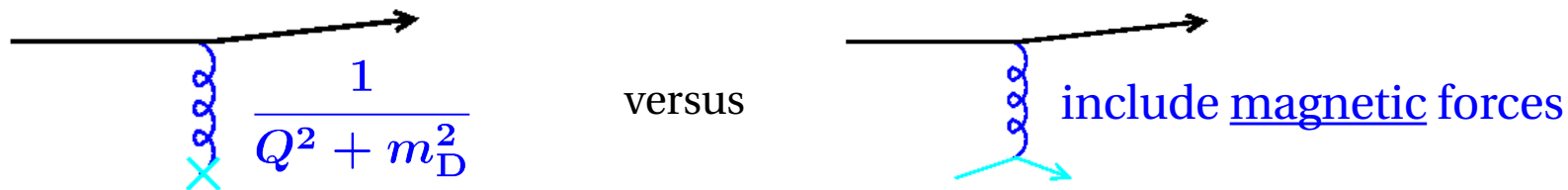
Does it only handle *soft* gluon bremsstrahlung?



versus

+

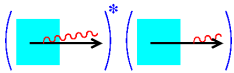
Does it assume “static” scatterers?



Does it include final-state Bose enhancement or Fermi blocking factors for plasma particles?

$n \frac{d\sigma_{\text{el}}}{d^2 q_{\perp}}$ 
versus
 $\frac{d\Gamma_{\text{el}}}{d^2 q_{\perp}} = \int dq_z \int d^3 p_2 \frac{d\sigma_{\text{el}}}{d^3 q} f(p_2) [1 \pm f(\vec{p}_2 - \vec{q})]$

Issues on this page are relevant if you want to get exactly the correct answer in the weak coupling limit.

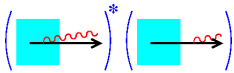
	thickness		non-uniform media?	$x$ values	non-static scatterers and $1 \pm f$ ?	exact for small $\alpha$ ?
Zakharov	any	yes	yes	any	no	no
BDMPS	any	yes	yes	any	no	no

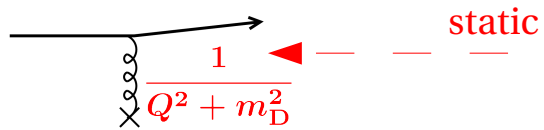
BDMPS ('96)

Zakharov ('96)





	thickness		non-uniform media?	$x$ values	non-static scatterers and $1 \pm f$ ?	exact for small $\alpha$ ?
Zakharov	any	yes	yes	any	no	no
BDMPS	any	yes	yes	any	no	no
BDMPS <sub>model</sub>	any	yes	yes	any	no	no



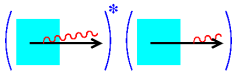
BDMPS ('96)

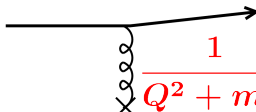
Zakharov ('96)

equivalence ('98)

a problem in non-Hermitian 2-D quantum mech.

$$H(t) = \frac{p_{\perp}^2}{2\mathcal{M}} - i\Gamma(x_{\perp}, t)$$

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Zakharov	any	yes	yes	any	no	no
BDMPS	any	yes	yes	any	no	no
BDMPS <sub>model</sub>	any	yes	yes	any	no	no
BDMPS <sub>HO</sub>	thick	yes	yes	any	no	no



static

1

$Q^2 + m_D^2$

BDMPS ('96)

Zakharov ('96)

equivalence ('98)

a problem in non-Hermitian 2-D quantum mech.

thick

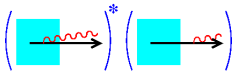
$\ln(E/T) \gg 1$

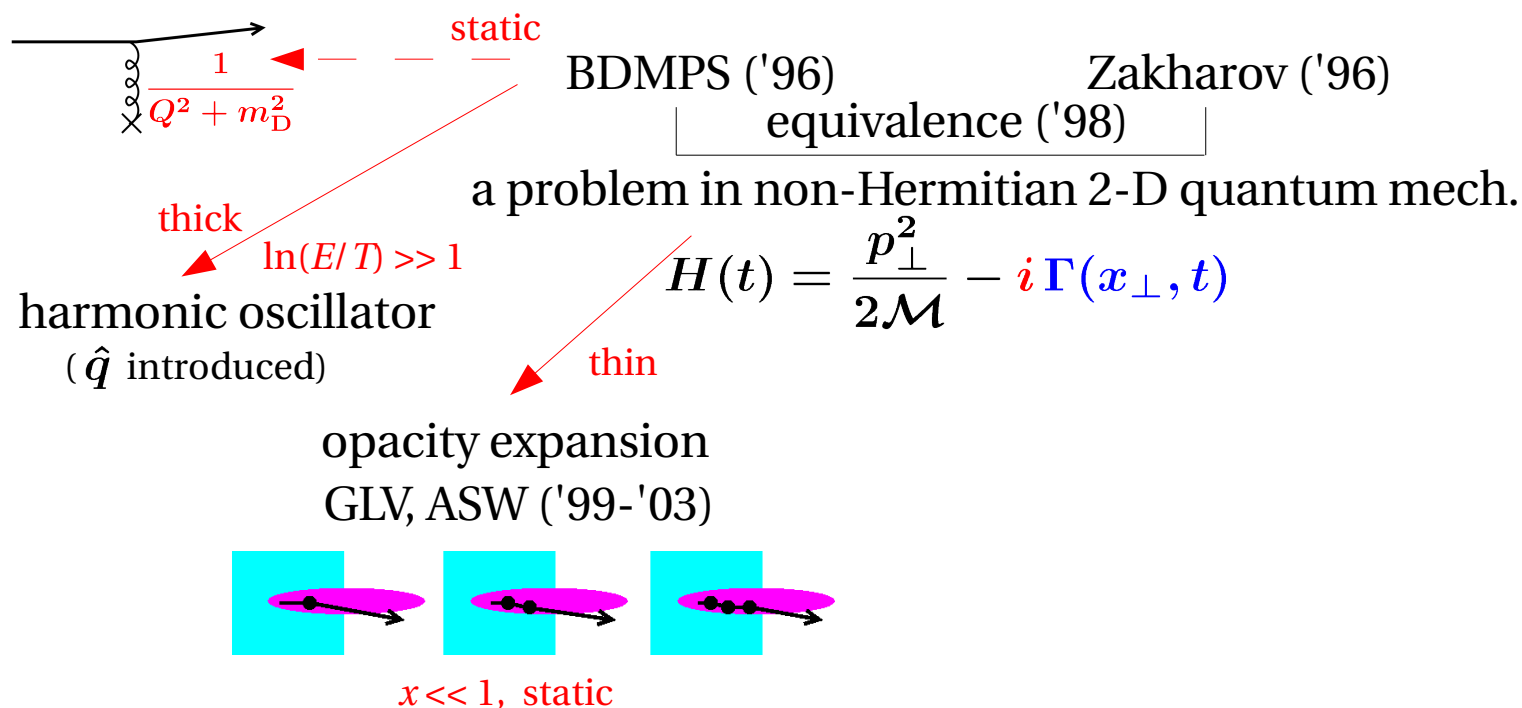
harmonic oscillator

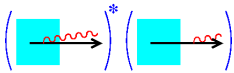
( $\hat{q}$  introduced)

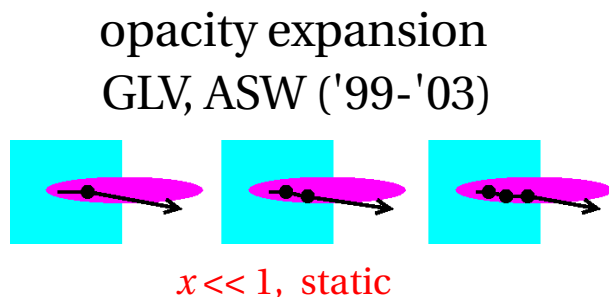
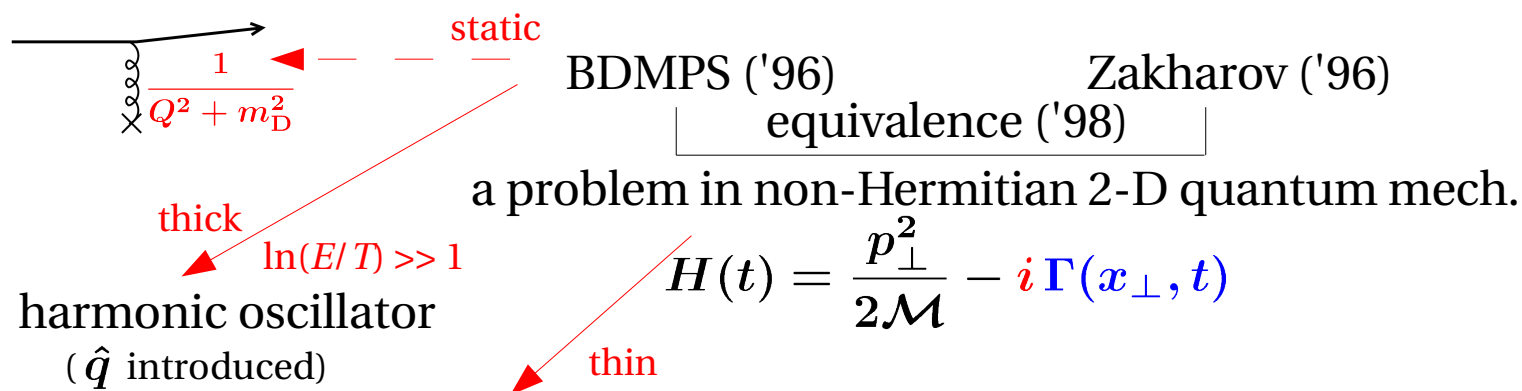
$H(t) = \frac{p_\perp^2}{2\mathcal{M}} - i\hat{q}x_\perp^2$

$H(t) = \frac{p_\perp^2}{2\mathcal{M}} - i\Gamma(x_\perp, t)$

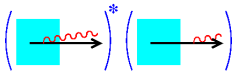
	thickness		non-uniform media?	$x$ values	non-static scatterers and $1 \pm f$ ?	exact for small $\alpha$ ?
Zakharov	any	yes	yes	any	no	no
BDMPS	any	yes	yes	any	no	no
BDMPS <sub>model</sub>	any	yes	yes	any	no	no
BDMPS <sub>HO</sub>	thick	yes	yes	any	no	no
ASW	any/thin	yes	yes	$x \ll 1$	no	no

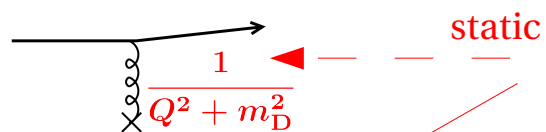


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AMY	"infinite"	N/A	uniform	any	yes	yes



AMY ('02-'03)  
infinite media  
exact for small  $\alpha$

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AMY	"infinite"	N/A	uniform	any	yes	yes
generalized BMDPS-Z	any	yes	yes	any	yes	yes



BDMPS ('96)

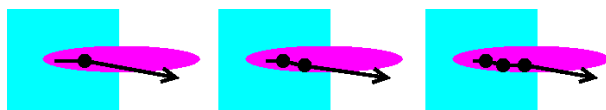
Zakharov ('96)

equivalence ('98)

a problem in non-Hermitian 2-D quantum mech.

thick  
 $\ln(E/T) \gg 1$   
 harmonic oscillator  
 ( $\hat{q}$  introduced)

thin  
 opacity expansion  
 GLV, ASW ('99-'03)



$x \ll 1$ , static

AMY ('02-'03)

infinite media  
 exact for small  $\alpha$

generalized BMDPS-Z  $\rightarrow$  AMY ('09)

$n d\sigma_{\text{el}} \rightarrow d\Gamma_{\text{el}}$



# Summary

- Weak coupling ain't simple at high temperature – lots of rich, complicated physics.
- The LPM effect is easy to understand qualitatively!
- There's a simple generalization of earlier formalisms for calculating the LPM effect in QCD that will yield exact results in the weak coupling limit if one simply uses weak-coupling results for the elastic scattering rate  $d\Gamma_{\text{el}}$ .

# Issues with weak coupling analysis

**Practical issue:** How big is the next-order correction in  $\alpha_s$ ?

$$\text{result} = (\text{leading order}) [ 1 + O(g) ]$$

How big can  $\alpha_s$  be before correction is 100% effect?

Example:  $d^2\Gamma_{\text{el}}/dq_{\perp}^2$

[Caron-Huot '09]



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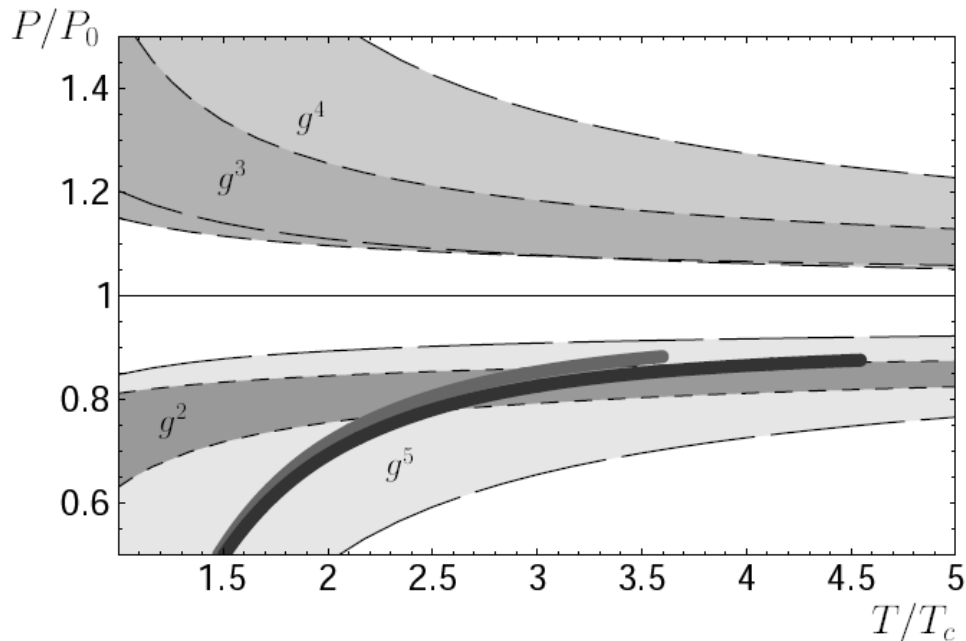
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Similar to long-standing problem with the QCD equation of state:



$g^2$  and  $g^3$  corrections the same size when

$$\alpha_s \sim 0.1 \sim \alpha_s(100 \text{ GeV})$$

Folks have tried various resummations of perturbation theory...

# Issues with weak coupling analysis

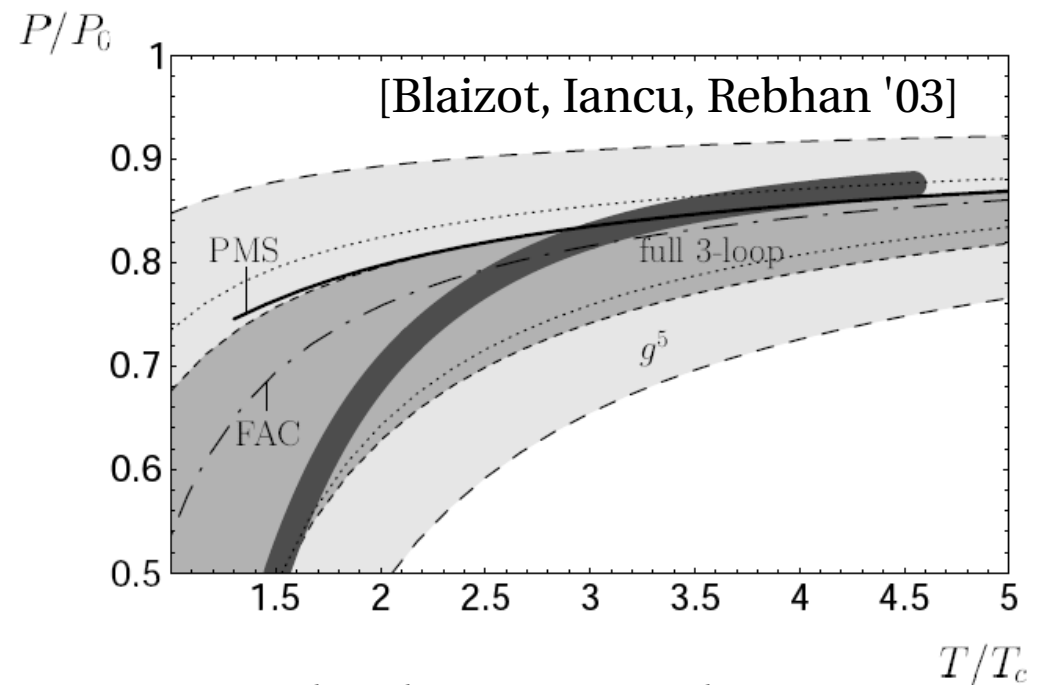
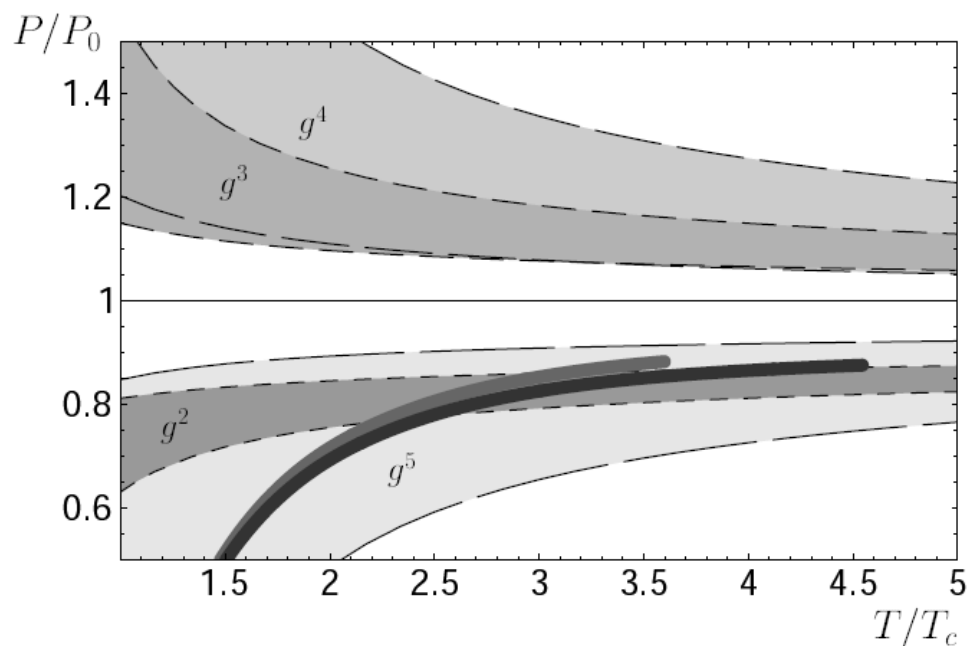
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Similar to long-standing problem with the QCD equation of state:



Not clear how to generalize to dynamics...

# Issues with weak coupling analysis

## Theoretical issue:

**Weak coupling**  $\alpha_s(T) \ll 1$

If LPM effect ignored:

$$\text{stopping distance} \propto \ln E$$

Actual result:

$$\text{stopping distance} \propto \left( \frac{E}{\ln E} \right)^{1/2}$$

**Strong coupling**  $\alpha_s \rightarrow \infty$  in large- $N_c$   $\mathcal{N}=4$  **SUSY** QCD

$$\text{stopping distance} \propto E^{1/3}$$

What's the first correction to the exponent for small $\alpha$ ?
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